

VARIANCE REDUCTION FOR MONTE CARLO CRITICALITY CALCULATIONS USING ADAPTIVE MULTILEVEL SPLITTING

Kévin Fröhlicher (IRSN)

Eric Dumonteil (CEA/IRFU)

Loïc Thulliez (CEA/IRFU)

Julien Taforeau (IRSN)

Mariya Brovchenko (IRSN)

Motivations

Can we tweak the Adaptive Multilevel Splitting (shielding) to compute a flux distribution in multiplicative media ?

What are the benefits of such method on variance, clustering, ... ?

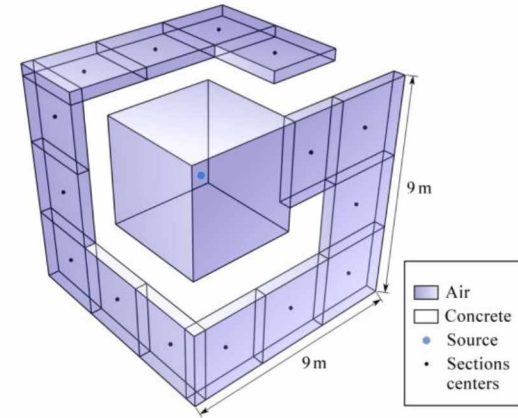
What further applications of the method in reactor physics might be possible?

Adaptive Multilevel Splitting (AMS) – State of the art

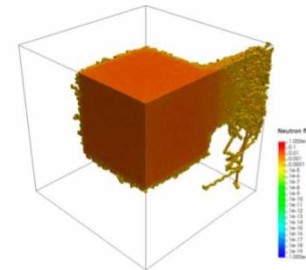


- Implemented at CEA for variance reduction for radiation protection calculations [1]
- The idea is to attract particles towards a target (detector) by **ranking** and **re-sampling** particle histories
- The method is **robust, easy to use** and has shown good results in shielding !

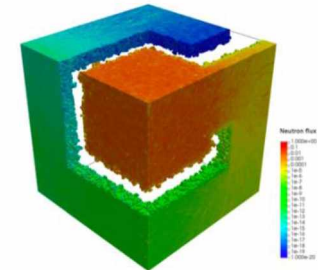
Can we adapt it to criticality ?



From [1] : H. Louvin (2017)



(a) Analog neutron flux



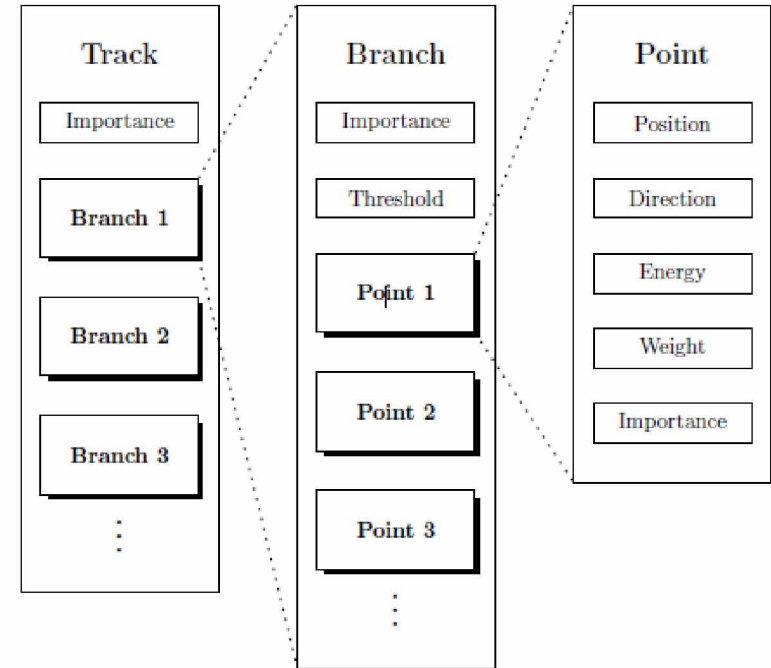
(b) AMS neutron flux

Attenuation $\sim 10^{20}$!

The AMS structure

- AMS track = neutron history
- AMS branch = one particle
- AMS point = collision, boundary crossing, ...

**Re-sampling tracks → re-sampling
“independent” neutron histories**



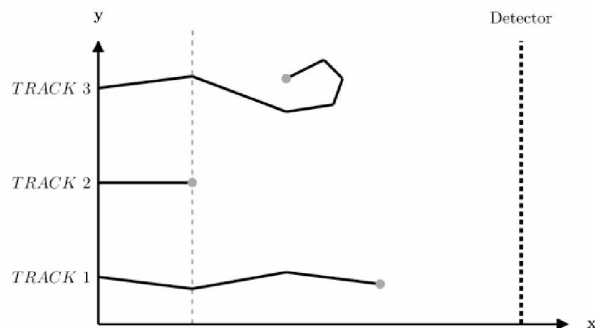
From [1] : H. Louvin (2017)

How does the AMS iterate

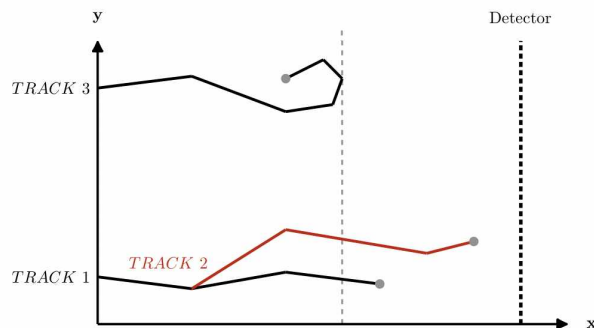
$N = 3$ tracks

$K = 1$ re-sampled at each iteration

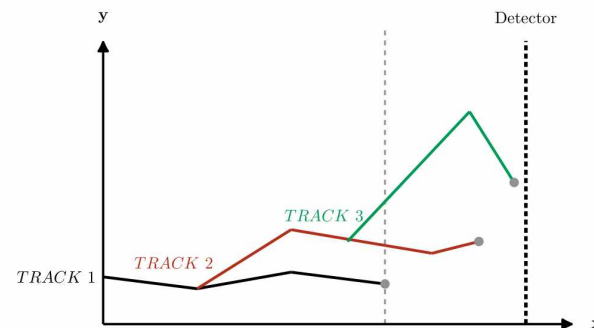
$$\text{Importance} \propto \frac{1}{\text{distance to the detector}}$$



Iteration 0 :
Analog transport of
all tracks



Iteration 1 :
Track 2 is deleted
and re-sampled



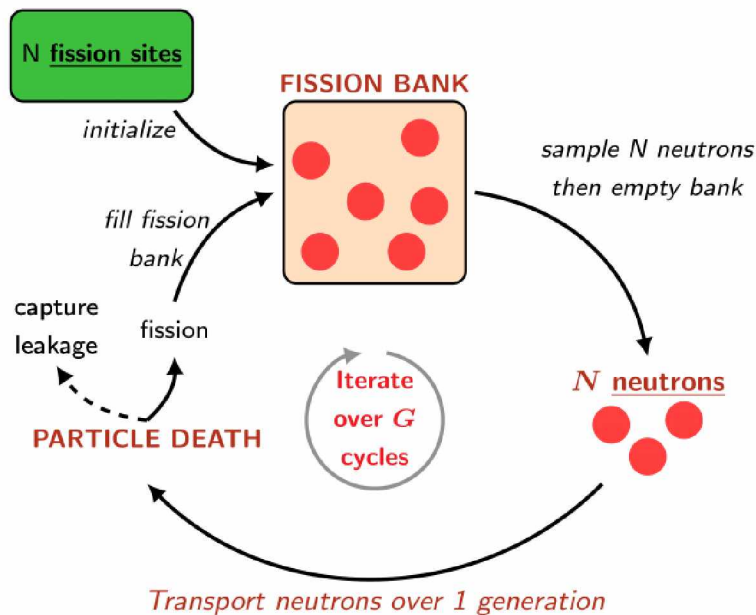
Iteration 2 :
Track 3 is deleted
and re-sampled

Final iteration : $\tilde{\phi} = \phi_{det} \times \alpha$

Probability to reach the detector : $\alpha = \left(1 - \frac{K}{N}\right)^L$

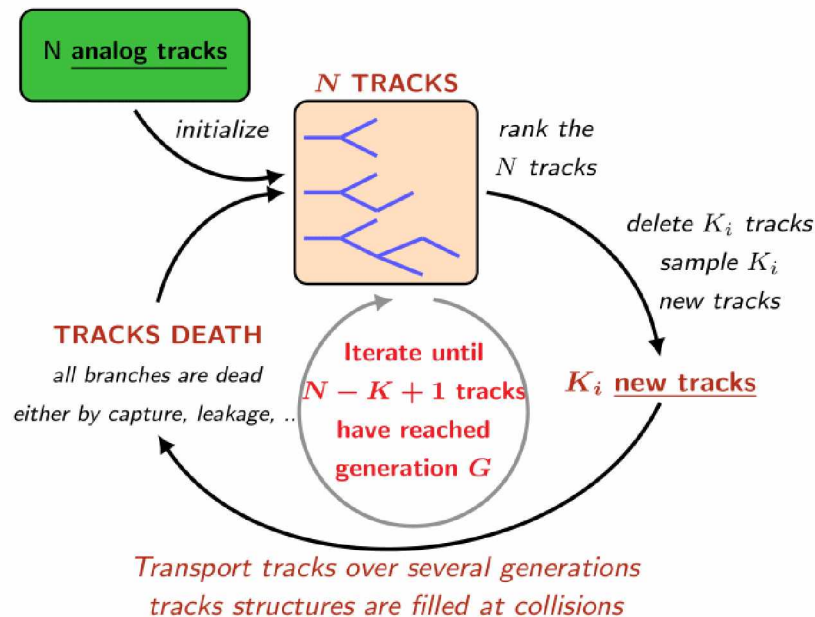
Application to criticality

Power iteration



AMS

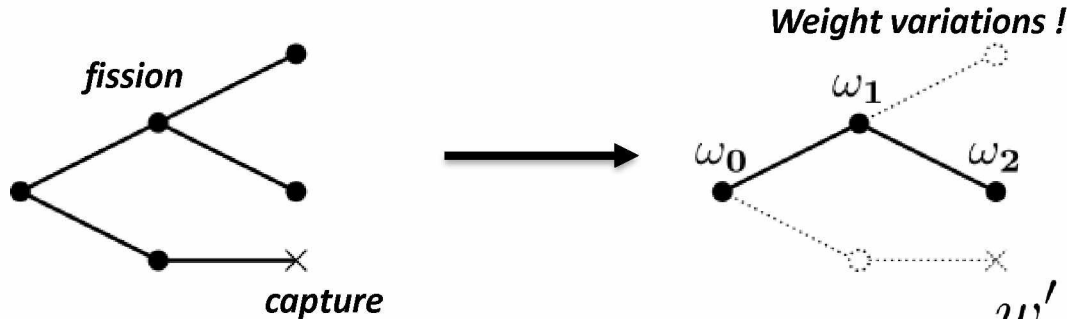
Importance $\propto$ generation



For AMS : criticality seen as an attenuation problem over successive generations

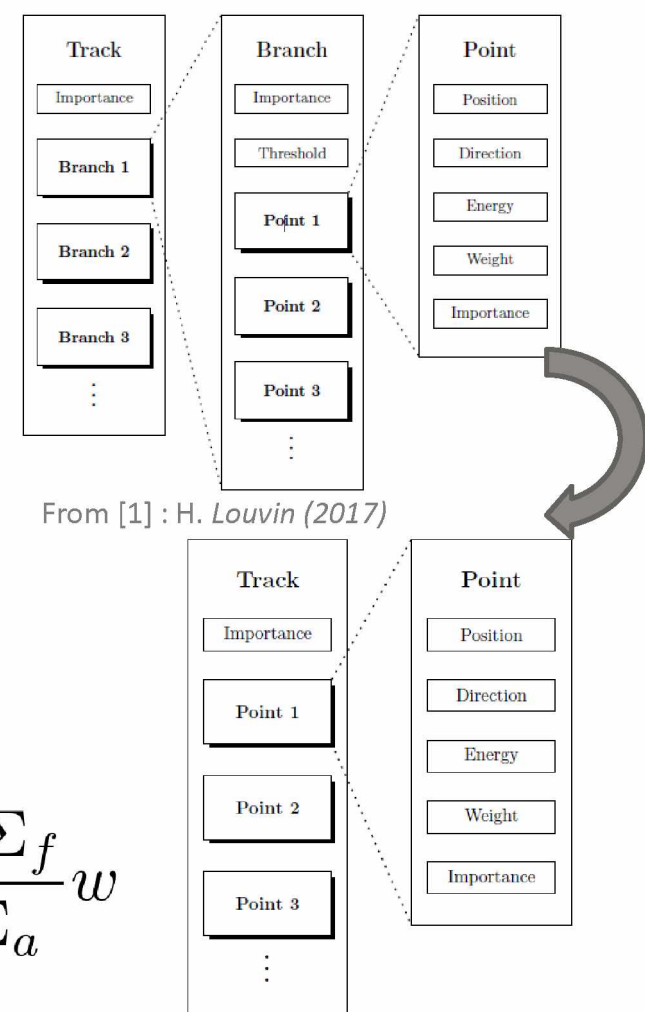
Handling multiplicative medium

- Possible to handle branching tracks [1], but...
- Increasing number of branches currently limits the use of AMS for **subcritical** systems
- We need a **numerically subcritical** system : **branchless collision** [2]



No more branches

$$w' = \frac{\bar{\nu} \Sigma_f}{\Sigma_a} w$$



Application to a one-dimensionnal bare slab

PROBLEM

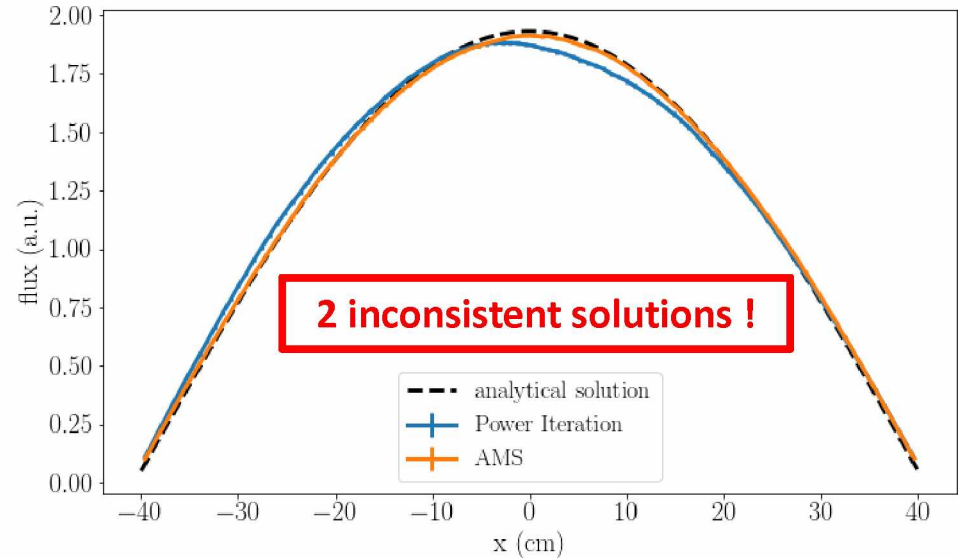
- 80 cm One-dimensionnal bare slab
- Leakage
- One-speed problem
- 1000 neutrons/cycle, 1000 cycles, 100 independent runs
- Flux $\pm 3\sigma$

AMS : $I = g + \frac{1}{1+|x|}$

Push neutrons
through generations

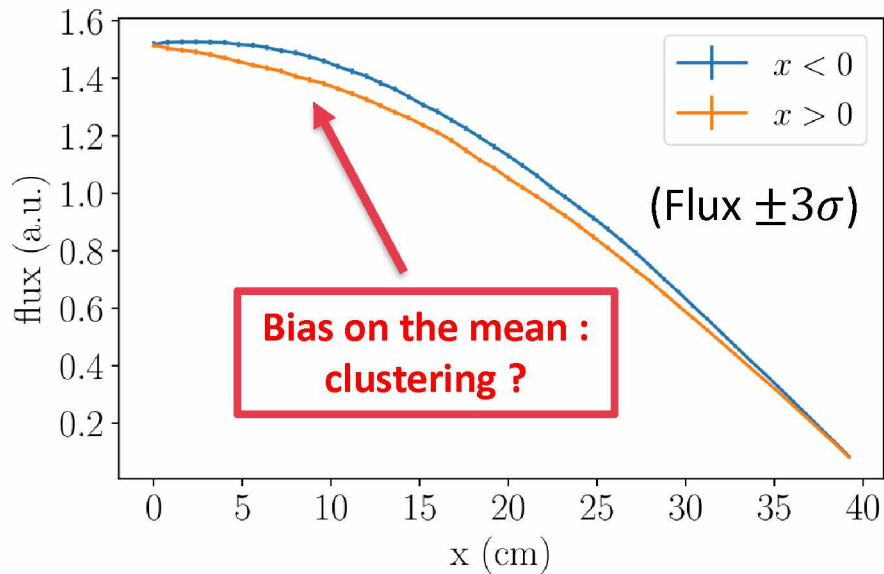
To avoid accretion
in one generation

FLUX

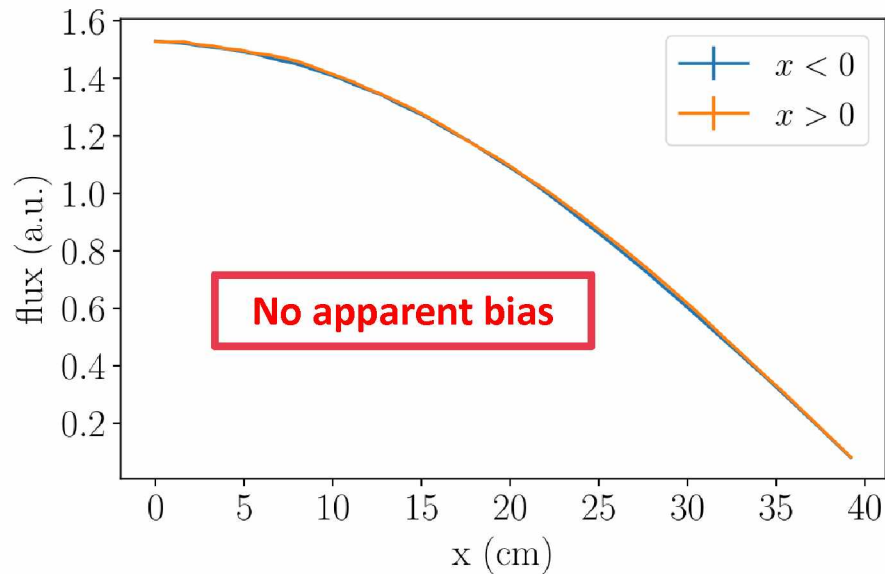


Homogeneous rod : symmetry of the solution

POWER ITERATION



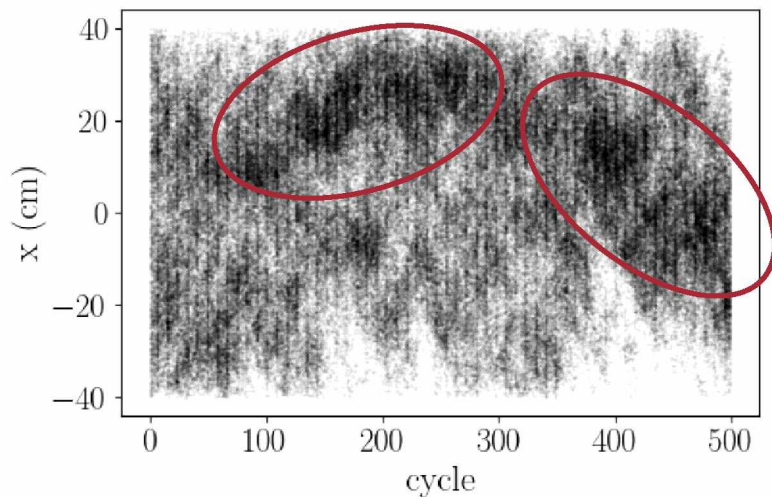
AMS + branchless



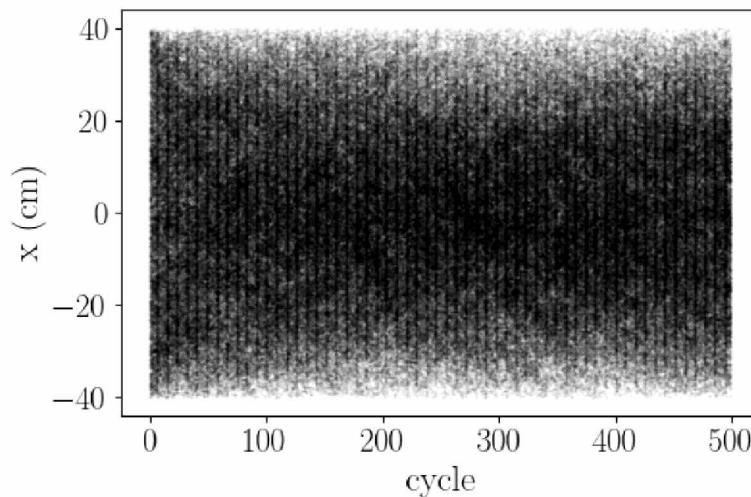
Formation of clusters in one simulation

Clusters are born from the imbalance between the birth (localized) and death (uniform) processes

POWER ITERATION



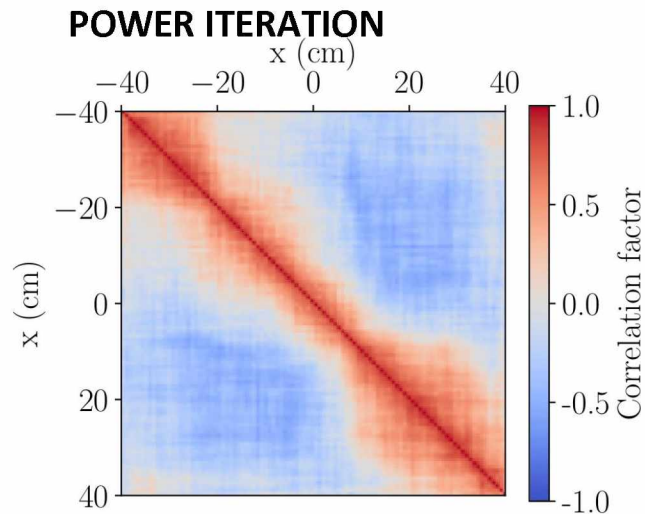
AMS + branchless



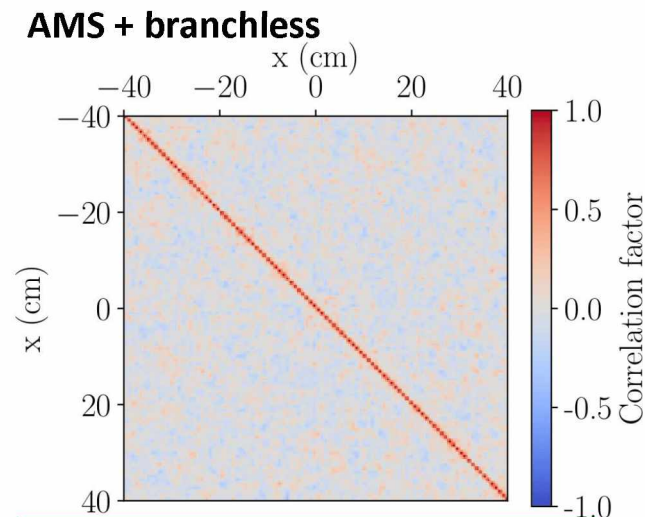
AMS + branchless : fewer visible clusters

Spatial correlations

In one generation...



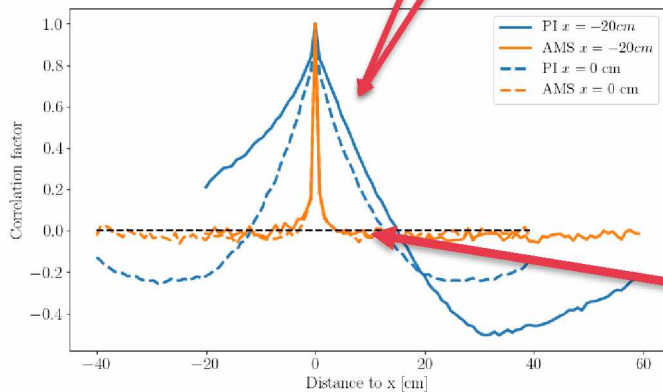
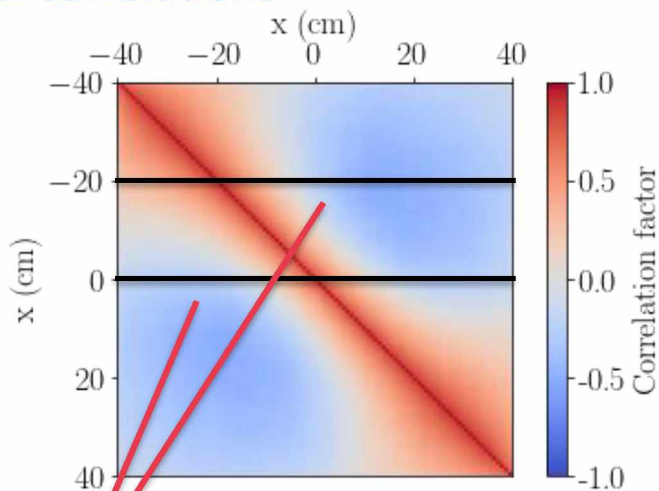
$$\rho_{i,j}(g) = \frac{\text{Cov} [\phi_g(x_i), \phi_g(x_j)]}{\sigma [\phi_g(x_i)] \sigma [\phi_g(x_j)]}$$



**Spatial bins are barely
correlated to each other !**

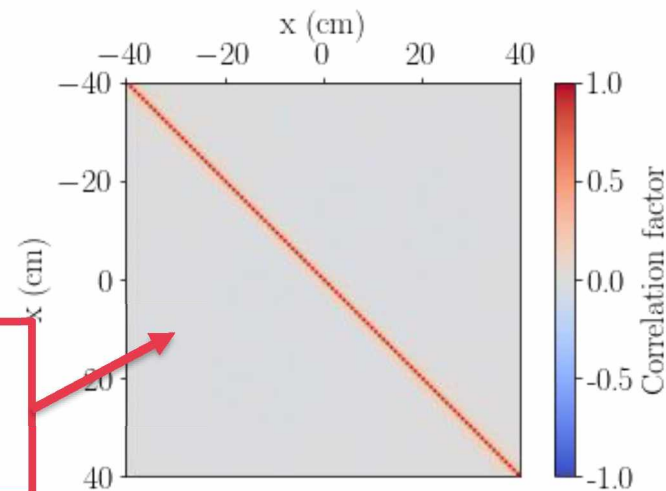
Mean spatial correlations

**POWER
ITERATION**



$$\rho_{i,j} = \langle \rho_{i,j}(g) \rangle_g$$

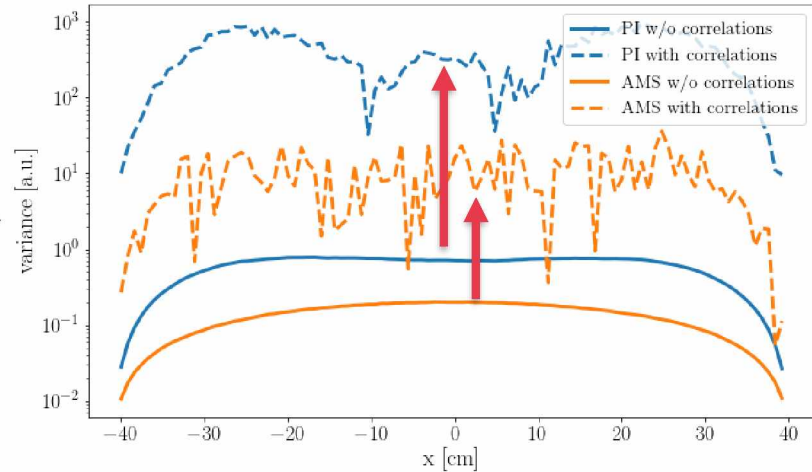
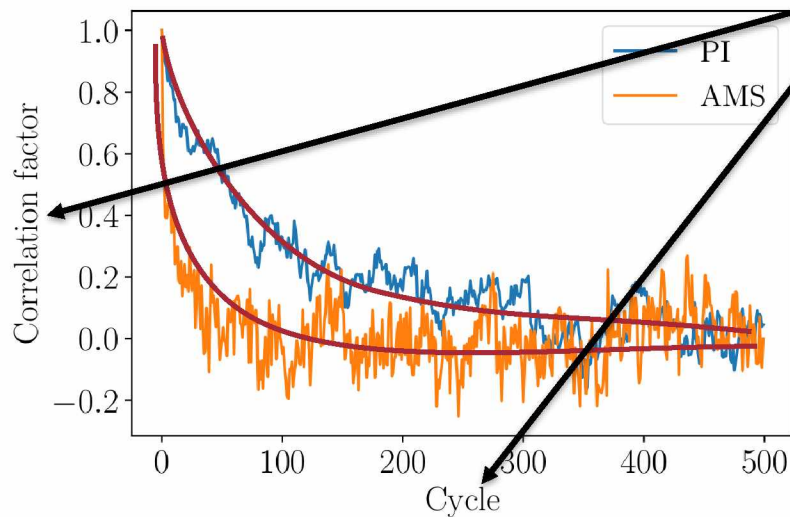
AMS + branchless



**Less chances for
clusters to appear
and develop**

Underestimation of the variance due to cycles correlations

Underestimation by a factor : $1 + 2 \sum_{k=1}^{n-1} \left(1 - \frac{k}{n}\right) \rho_k$



AMS + branchless : cycle-to-cycle correlations die more quickly



Real variance is less underestimated by AMS+branchless

Conclusion

Can we tweak the Adaptive Multilevel Splitting (shielding) to do criticality calculations ?

What are the benefits of such method on variance, clustering, ... compared to the power iteration ?

Reduces bias on the flux (clustering due spatial correlations)

Reduces variance (observed and real – cycle correlations)

What further applications of the method in reactor physics might be possible?

Criticality (variance reduction, rod depletion [3,4], ...)

Kinetics (population control and/or variance reduction)

**Thank you for
your attention**

References

- [1] Louvin, H. (2017). *Development of an adaptive variance reduction technique for Monte Carlo particle transport* (Doctoral dissertation, Université Paris-Saclay).
- [2] Sjenitzer, B. L., & Hoogenboom, J. E. (2011). Variance reduction for fixed-source Monte Carlo calculations in multiplying systems by improving chain-length statistics. *Annals of Nuclear Energy*, 38(10), 2195-2203.
- [3] Cosgrove, P., Shwageraus, E., & Parks, G. T. (2020). Neutron clustering as a driver of Monte Carlo burn-up instability. *Annals of Nuclear Energy*, 137, 106991.
- [4] Brovchenko, M., Kodratoff, N., Importance of the Axial Burn-Up Modelling for Fuel Assembly Nozzles Activation Calculation, *PHYSOR 2022*

Appendix

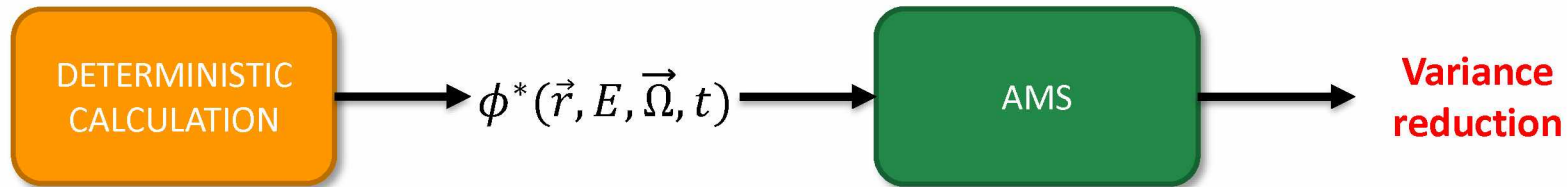
How to rank the tracks ?

- Particle importance : probability to contribute to the score

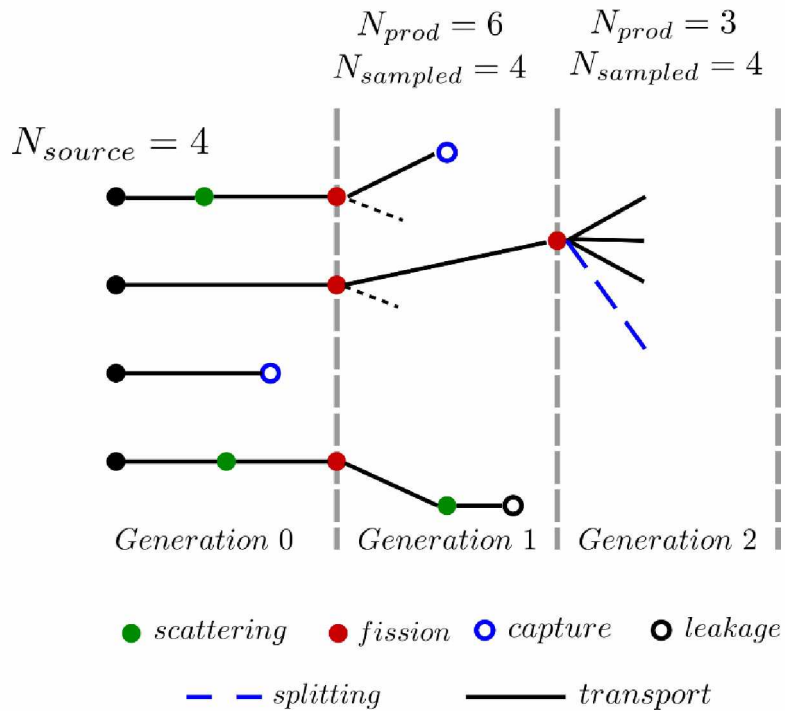
↳ $\text{importance in } \vec{r}, E, \vec{\Omega} \text{ at } t = \phi^*(\vec{r}, E, \vec{\Omega}, t)$

- Problem : adjoint calculation at least as difficult as the forward problem

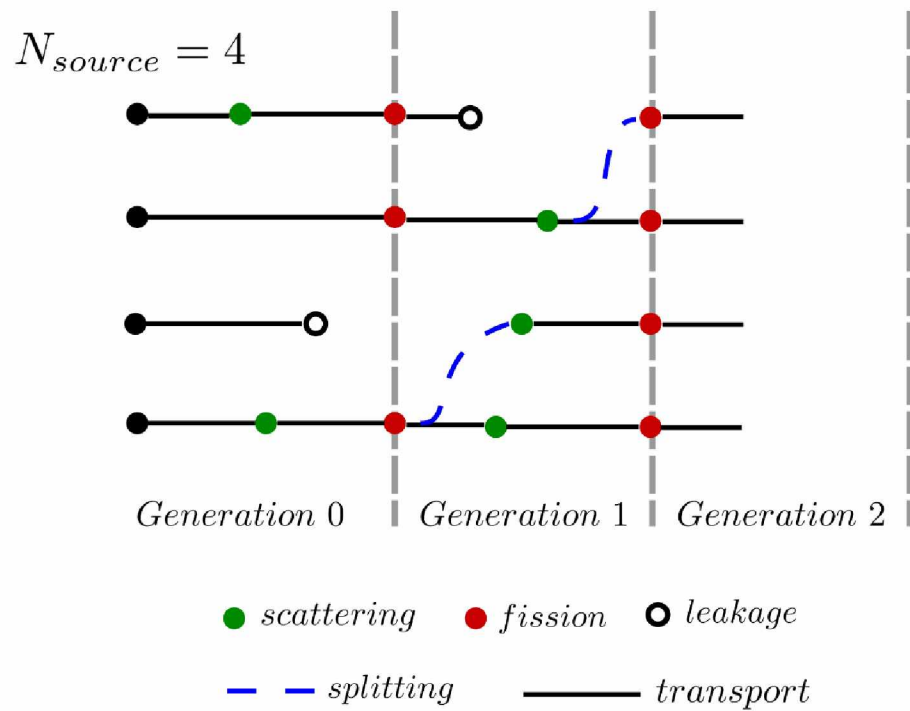
↳ **Even a rough estimation of ϕ^* could accelerate Monte Carlo !**



AMS + branchless VS Power Iteration



Power Iteration

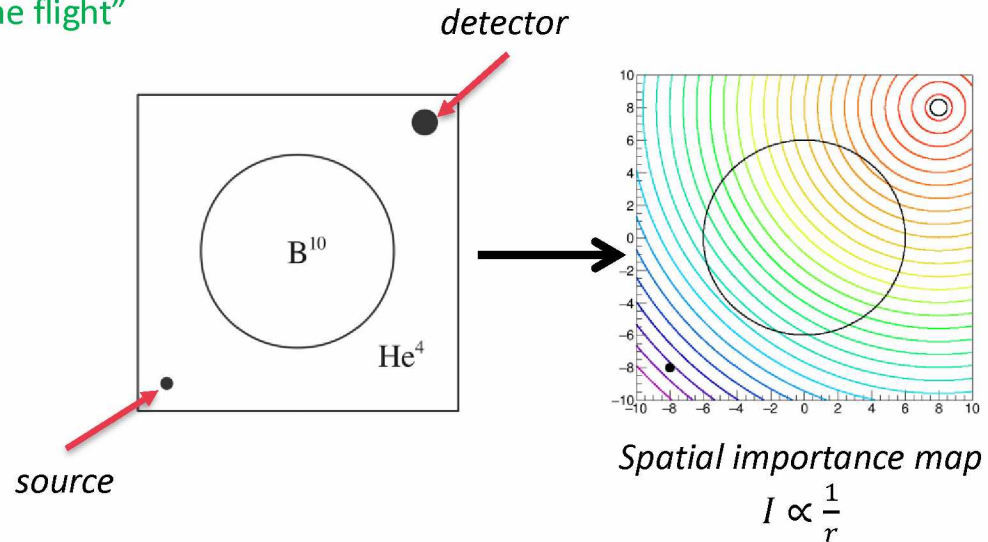


AMS with branchless collision

AMS for fixed source problems

AMS for shielding applications

- ❑ AMS = **Adaptive** Multilevel Splitting
- ❑ Originates from applied mathematics (Cerou et al, 2007)
- ❑ Applied to particle transport (Louvin et al, 2017)
- ❑ Defines the **splitting threshold** “on the flight”
- ❑ By **using an importance map**



AMS for fixed source problems : the algorithm

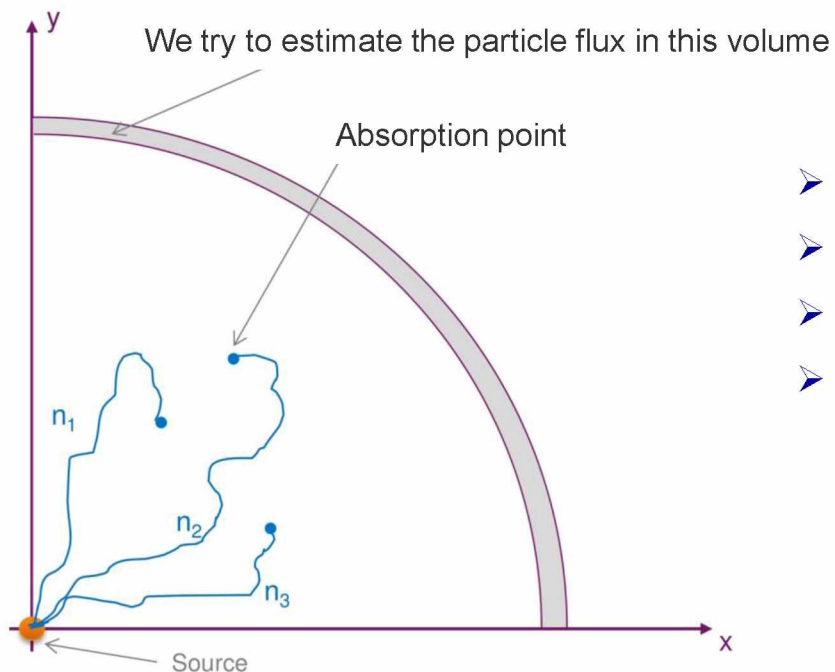
❑ Stopping criterium:

- When $n-k$ tracks have reached the “detector”, the algorithm stops
- The number of iteration corresponding to reach that criterium is N
- Each neutron is assigned a **statistical weight** α being:

$$\alpha = \left(1 - \frac{k}{n}\right)^N$$

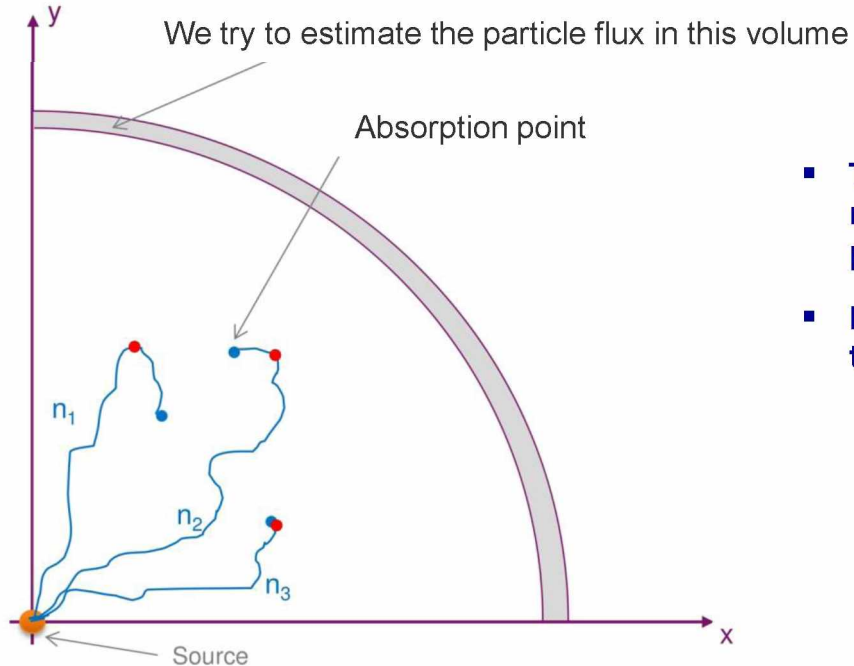
- ### ❑ An unbiased estimator of the flux is calculated “as usual” using the tracks of the last iteration

AMS for fixed source problems : the algorithm



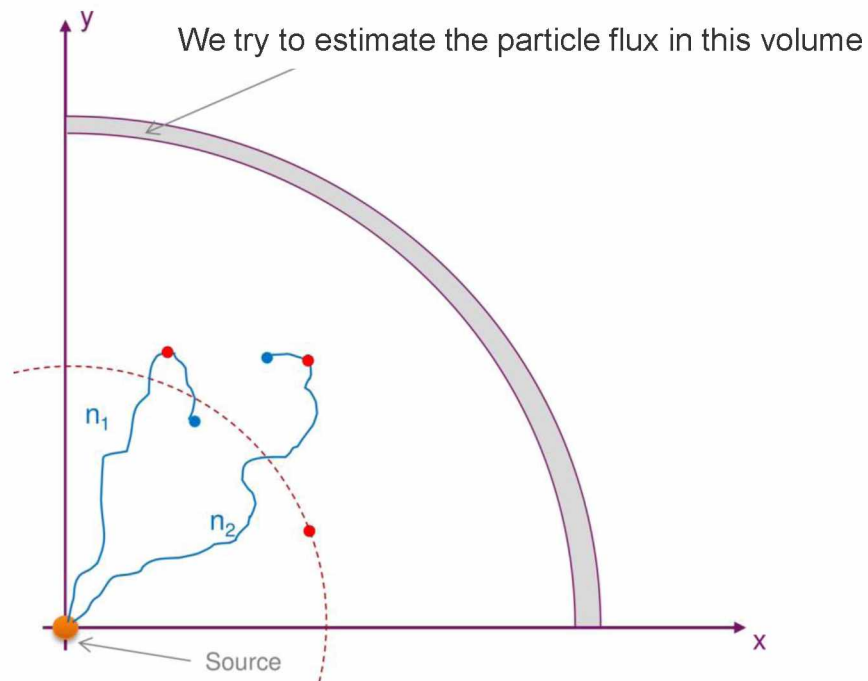
- $n=3$ initial particles
- $k=1$ track resampled
- importance = distance to the source
- Target: spherical shell (purple)

AMS for fixed source problems : the algorithm



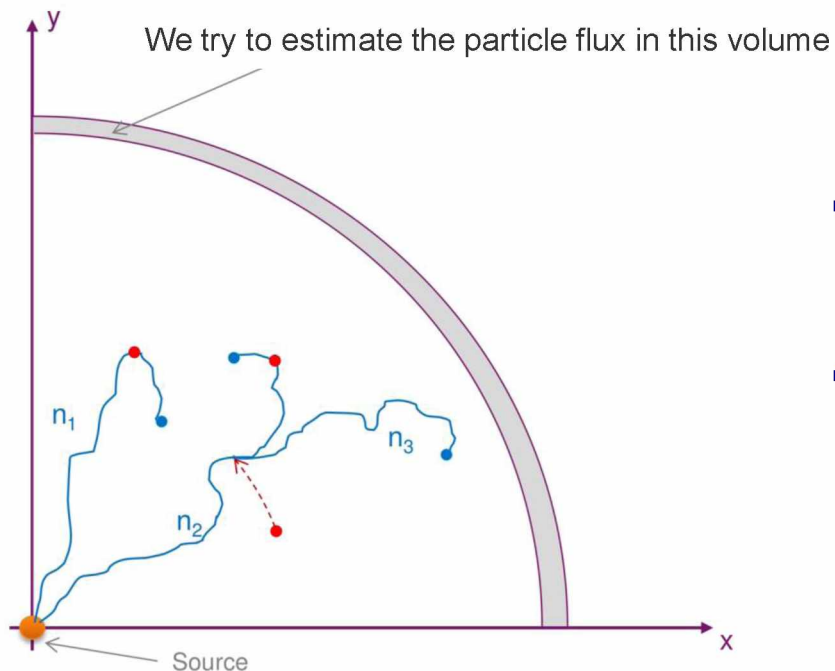
- The importance function is the maximal distance reached by the particle (red points)
- In this case the neutron tracks with the lowest score is n_3

AMS for fixed source problems : the algorithm



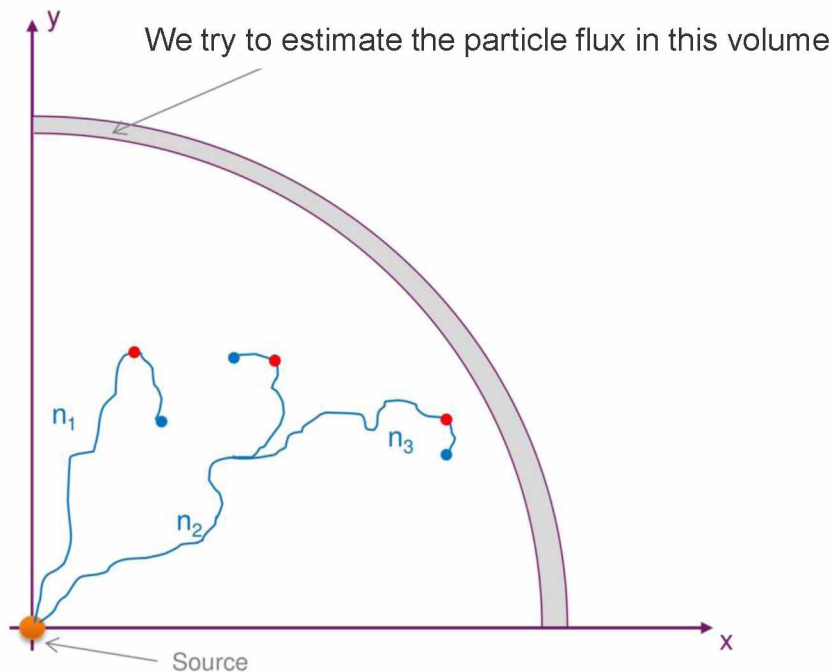
- The tracks associated to particle 3 is deleted
- The maximum score of this track is stored and defines the first splitting level

AMS for fixed source problems : the algorithm



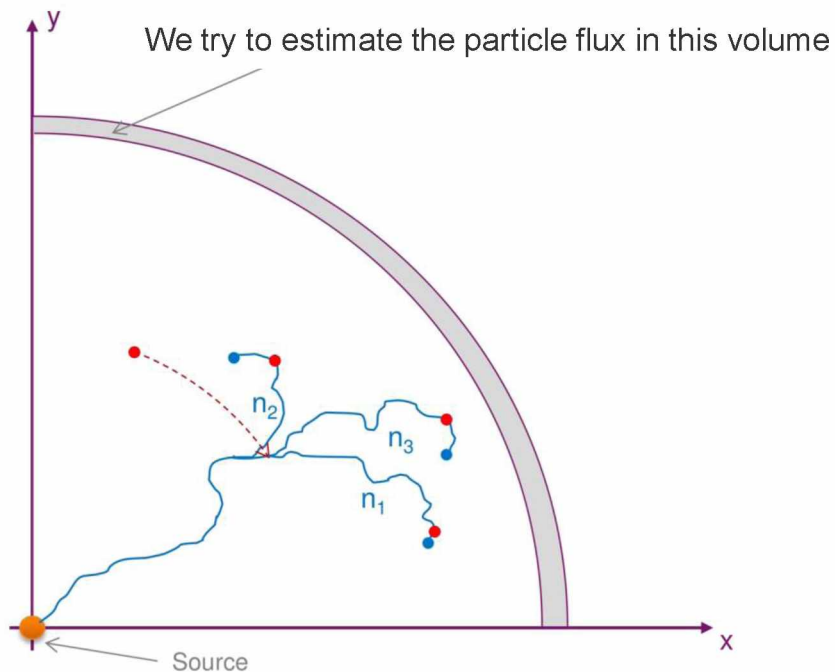
- Track number 2 is randomly sampled for the splitting
- A new particle is simulated from this splitting point until its absorption

AMS for fixed source problems : the algorithm



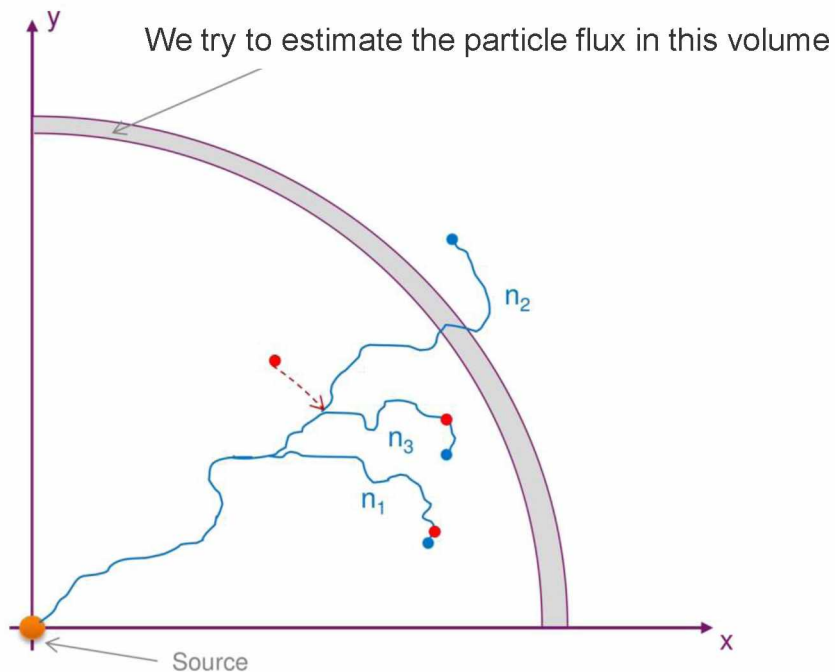
- The score of this new tracks n_3 is calculated
- The first iteration is over
- The stopping criterium is not meet: the iteration process goes on

AMS for fixed source problems : the algorithm



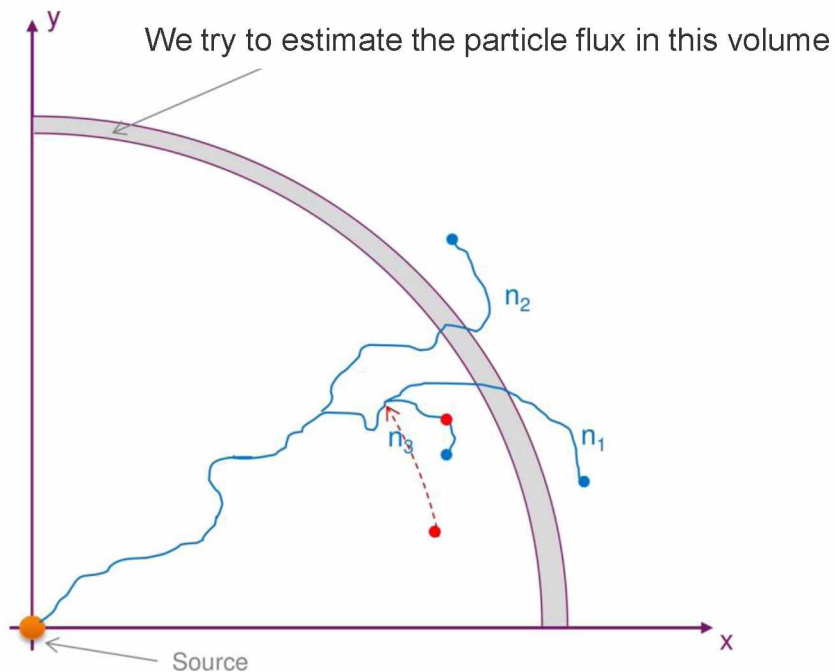
■ Iteration 2

AMS for fixed source problems : the algorithm



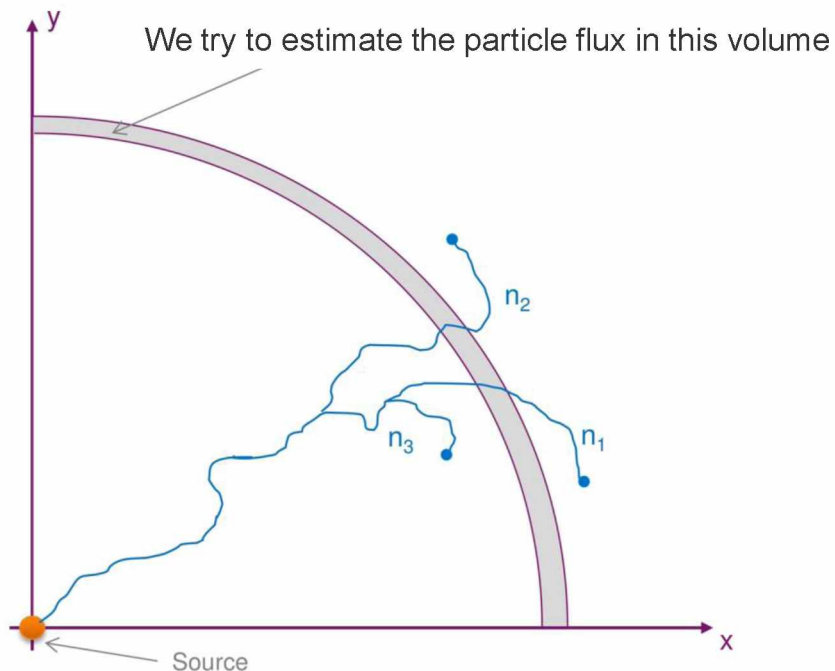
■ Iteration 3

AMS for fixed source problems : the algorithm



■ Iteration 4

AMS for fixed source problems : the algorithm



- Iteration 4
- $n-k$ particles have reached the target, the algorithm stops
- The statistical weight of **all** particles is :

$$\alpha = \left(1 - \frac{1}{3}\right)^4$$